

Theoretical review on systematic uncertainties for CP violation searches

Pilar Coloma



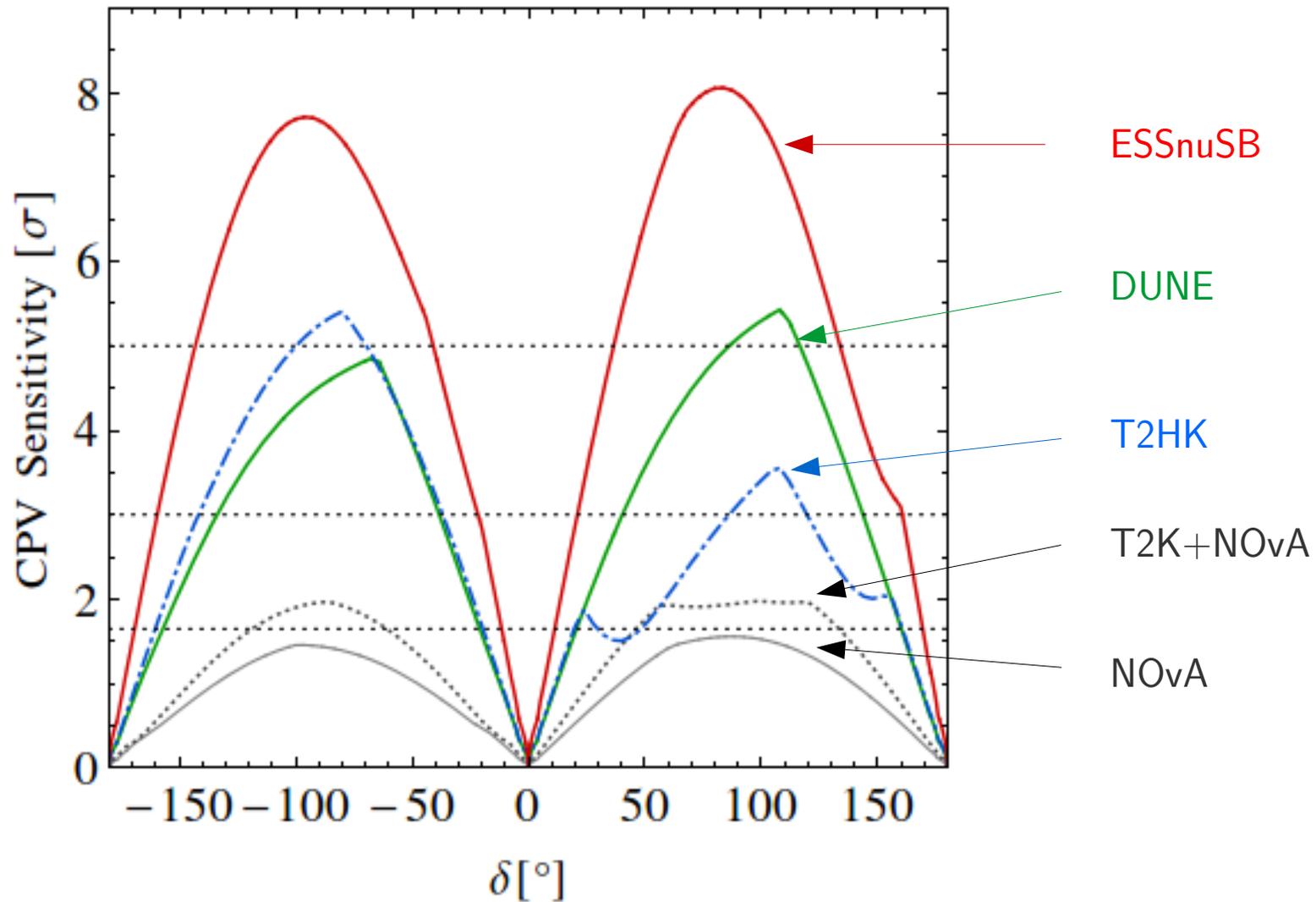
NNN'15

Stony Brook University, Oct 29th, 2015

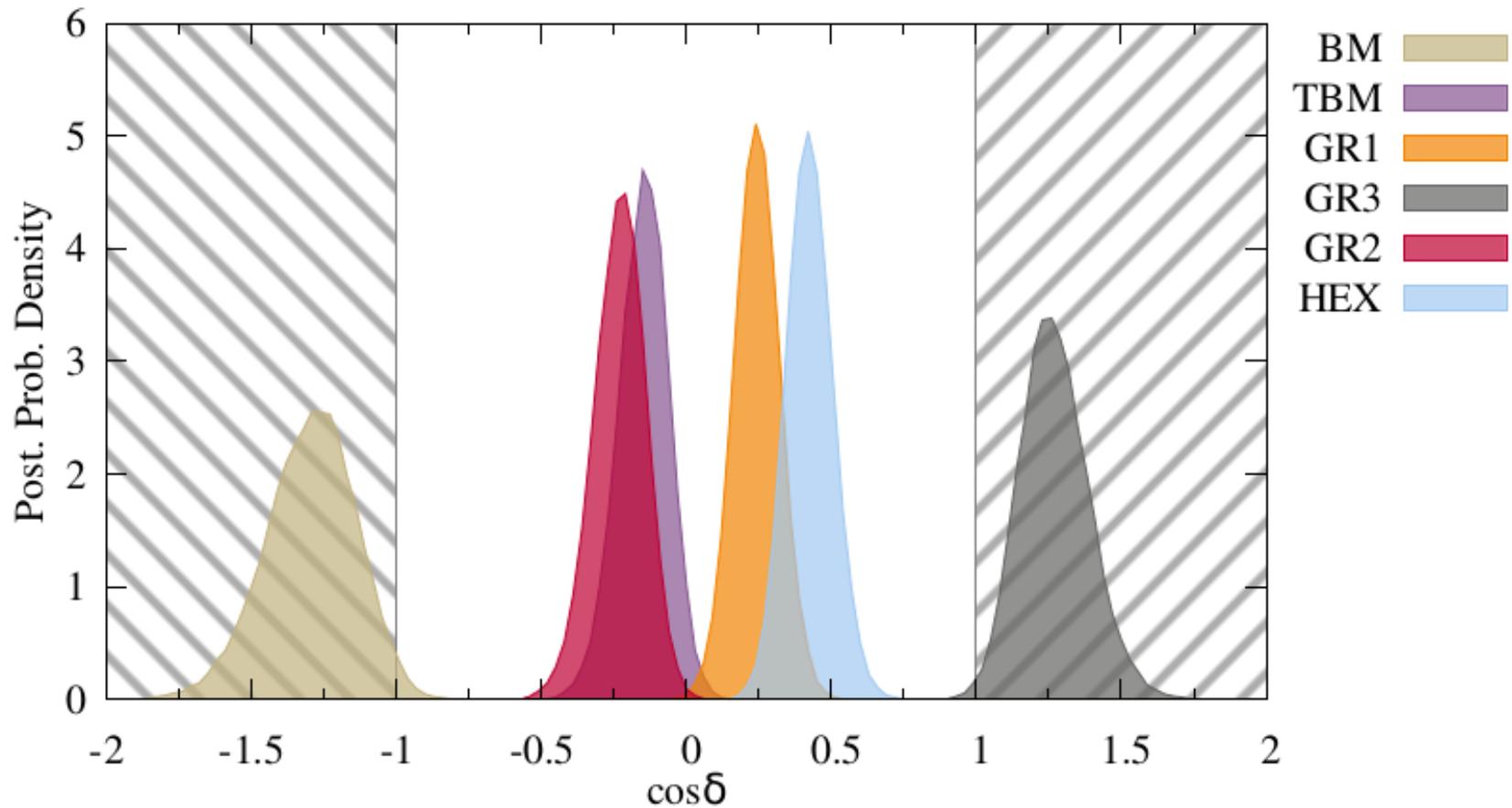
Outline

- 1) Introduction
- 2) Normalization uncertainties
- 3) Shape uncertainties
- 4) Summary

Proposed experiments for CPV searches



Why precision?



Ballett, King, Luhn, Pascoli, Schmidt, 1410.7573 [hep-ph]
(see also, e.g., Girardi et al, 1410.8056, Meloni, 1308.4578)

CP violation measurements

The **golden channel** in neutrino oscillations is:

$$P_{e\mu}^{\pm} \equiv P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu) = \overset{\text{atmospheric}}{s_{23}^2 \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta_{31} L}{2} \right)} + \overset{\text{solar}}{c_{23}^2 \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta_{21} L}{2} \right)} + \tilde{J} \cos \left(\pm\delta - \frac{\Delta_{31} L}{2} \right) \sin \left(\frac{\Delta_{21} L}{2} \right) \sin \left(\frac{\Delta_{31} L}{2} \right),$$

CP - violating interference

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{2E}$$

Cervera et al., hep-ph/0002108

CP violation measurements

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$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{2E}$$

Different for neutrinos and antineutrinos

Narrow band beams exploit this

CP violation measurements

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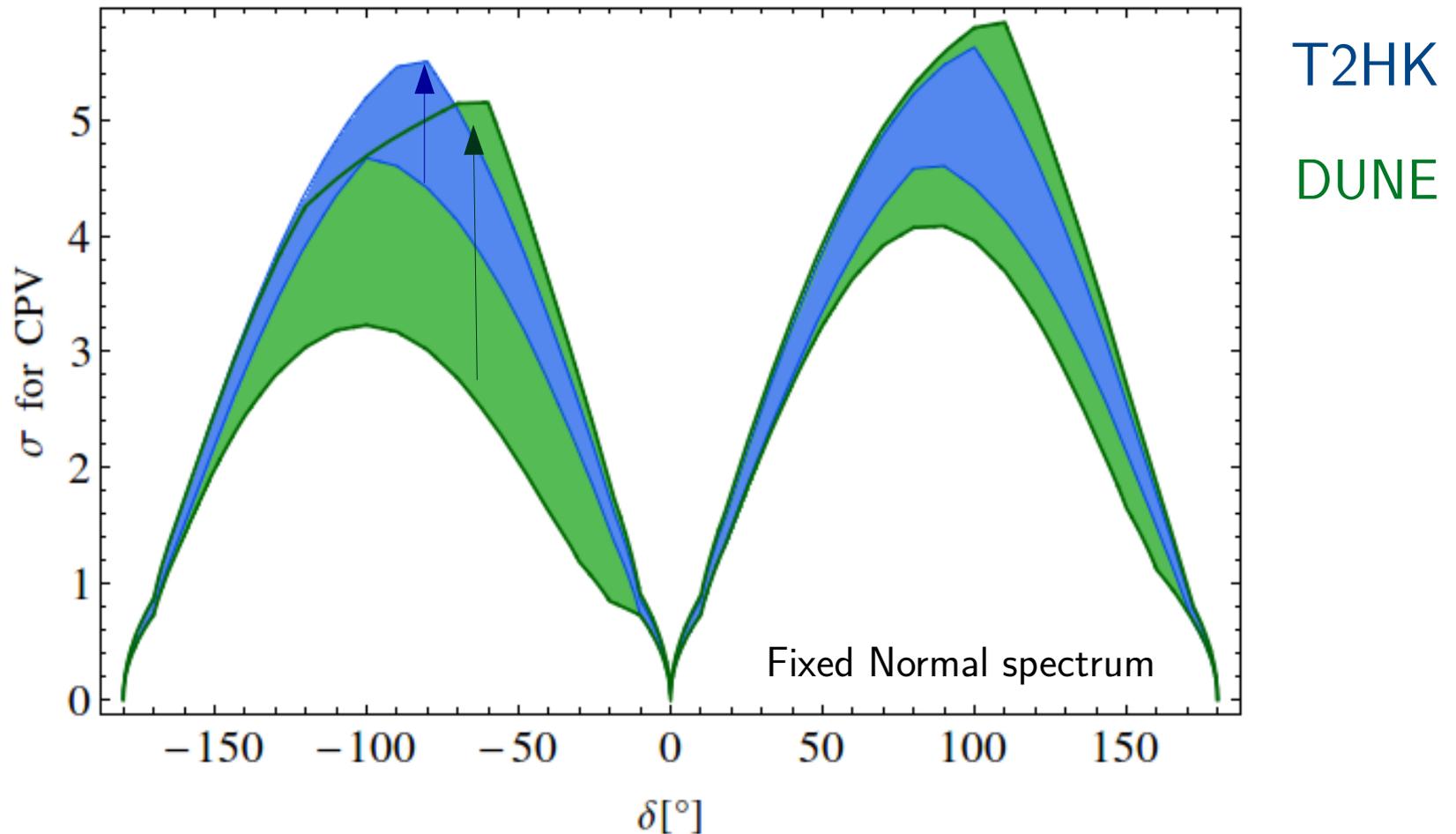
For neutrinos,
there is non-trivial
energy dependence
too

Wide band
beams exploit
this



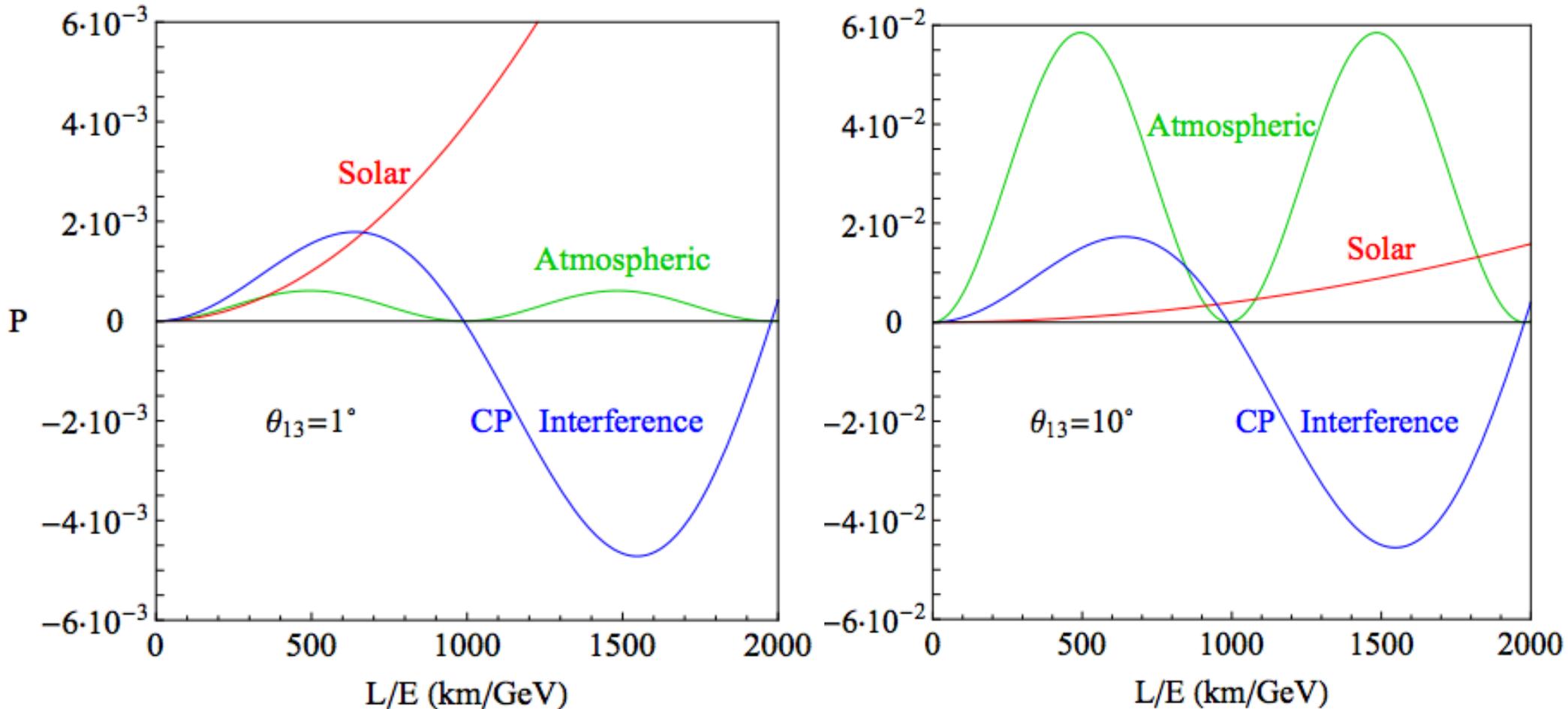
CP violation measurements

An example of relative importance of energy resolution:



Normalization uncertainties

Impact of systematics on CPV



Coloma and Fernandez-Martinez, 1110.4583 [hep-ph]

See also Marciano, hep-ph/0108181, Hagiwara et al, hep-ph/0607255 and Meregaglia, Rubbia, 0801.4035

Near/Far cancellation?

$$n_{\alpha \rightarrow \beta}(L, E) \sim \frac{1}{L^2} \epsilon_{\beta}(E) \times \sigma_{\beta}(E) \times \phi_{\alpha}(E) \times P_{\alpha\beta}(L, E)$$

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At reactor experiments, the cancellation of systematics between near/far detectors is very effective:

$$\frac{n_{ee}^{FD}}{n_{ee}^{ND}} \sim \frac{L_{ND}^2}{L_{FD}^2} \frac{\epsilon_e \sigma_e \phi_e}{\cancel{\epsilon_e \sigma_e \phi_e}} P_{ee}$$

Near/Far cancellation?

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At Daya Bay this works extremely well!

$$\sin^2 2\theta_{13} = 0.084^{+0.005}_{-0.005}$$

(Seminar given by Xin Qian at Fermilab, on Oct 15)

Near/Far cancellation?

- For CP violation searches, we need an appearance experiment
- An ideal near detector can be used to predict some backgrounds:

$$n_{\nu_e}^{FD,bg} \sim n_{\nu_e}^{ND,bg} \frac{L_{ND}^2}{L_{FD}^2} \frac{V_{FD}}{V_{ND}}$$

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- An ideal near detector can be used to predict some backgrounds:

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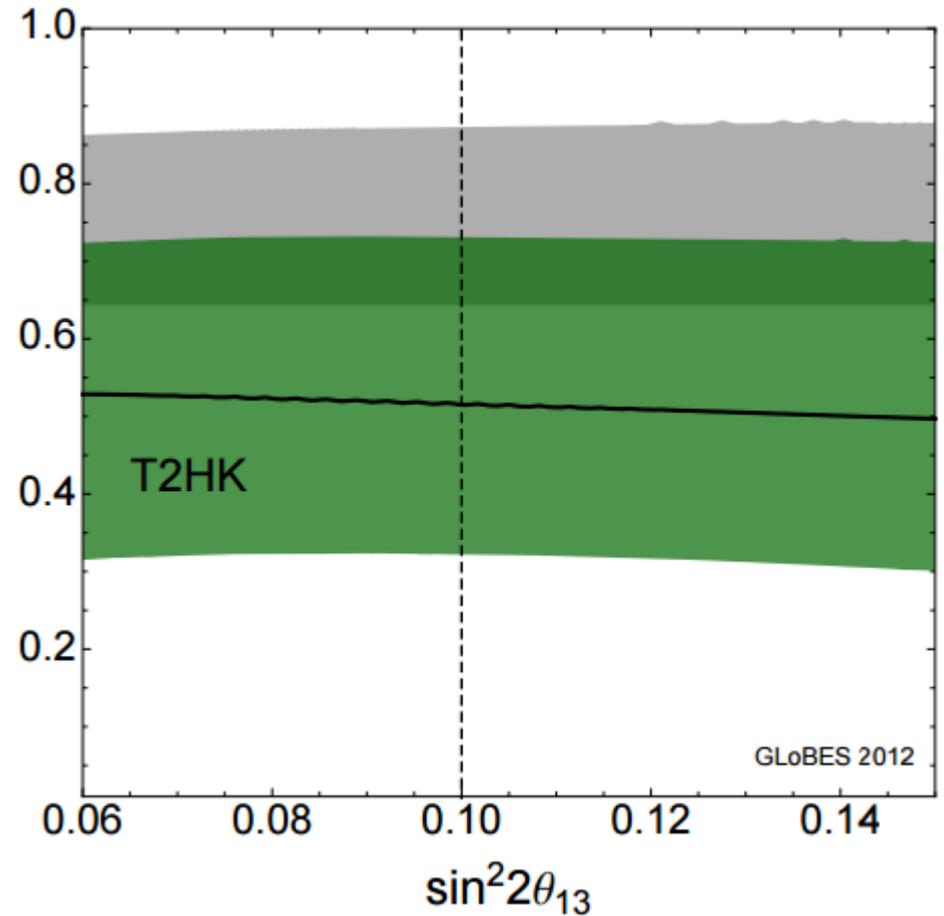
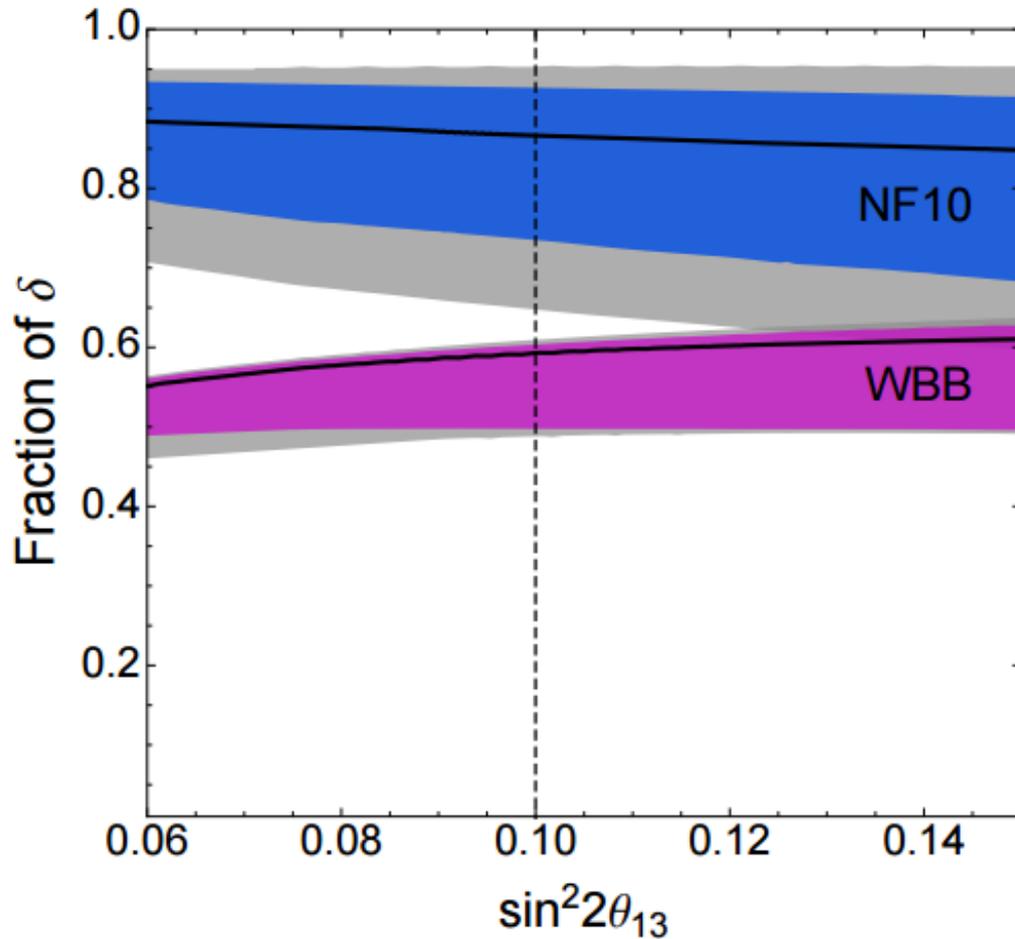
- However, a similar extrapolation for the signal will not work that well:

$$n_{\nu_e}^{FD,sig} \sim n_{\nu_\mu}^{ND,sig} \frac{L_{ND}^2}{L_{FD}^2} \frac{V_{FD}}{V_{ND}} \left(\frac{\tilde{\sigma}_{\nu_e}}{\tilde{\sigma}_{\nu_\mu}} \right) \times P(\nu_\mu \rightarrow \nu_e)$$

Huber, Mezzetto and Schwetz, 0711.2950 [hep-ph]

Coloma, Huber, Kopp, Winter, 1209.5973 [hep-ph]

Impact of normalization uncertainties



Coloma, Huber, Kopp, Winter, 1209.5973 [hep-ph]
(See also Huber, Mezzetto and Schwetz, 0711.2950 [hep-ph])

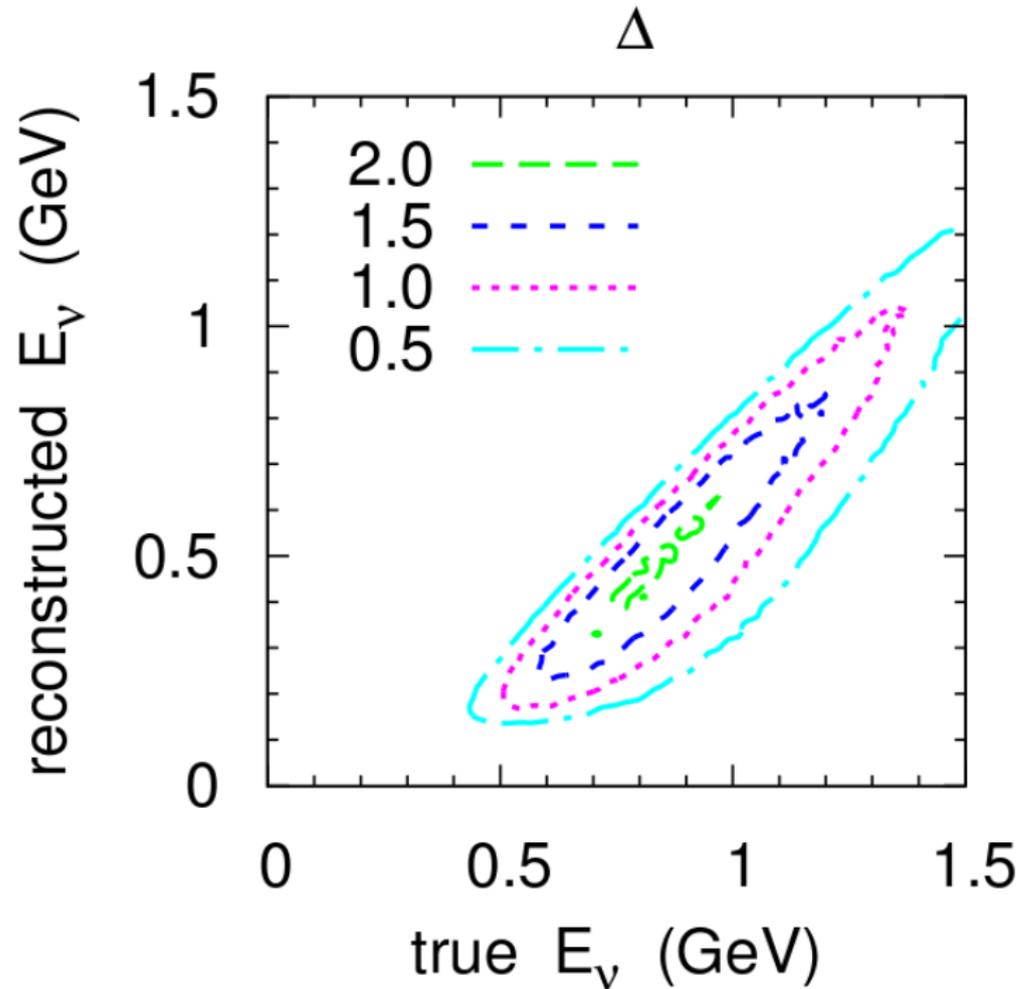
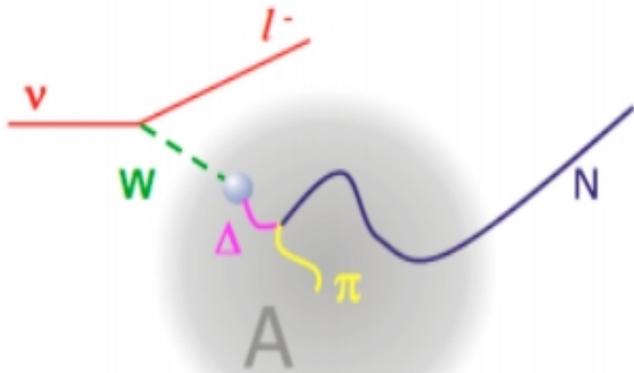
Possible ways out

- Correlations can help to reduce impact of systematics:
 - the far detector can act as a “near detector”
- Possible ways to reduce the effect of normalization uncertainties:
 - measure final flavor cross sections at a near detector (intrinsic contamination).
 - put theoretical constraints on ratios between cross sections for different flavors
- Caveats: near/far flux extrapolation is tricky; near/far detectors may not be identical, etc

Shape uncertainties

Energy reconstruction effects

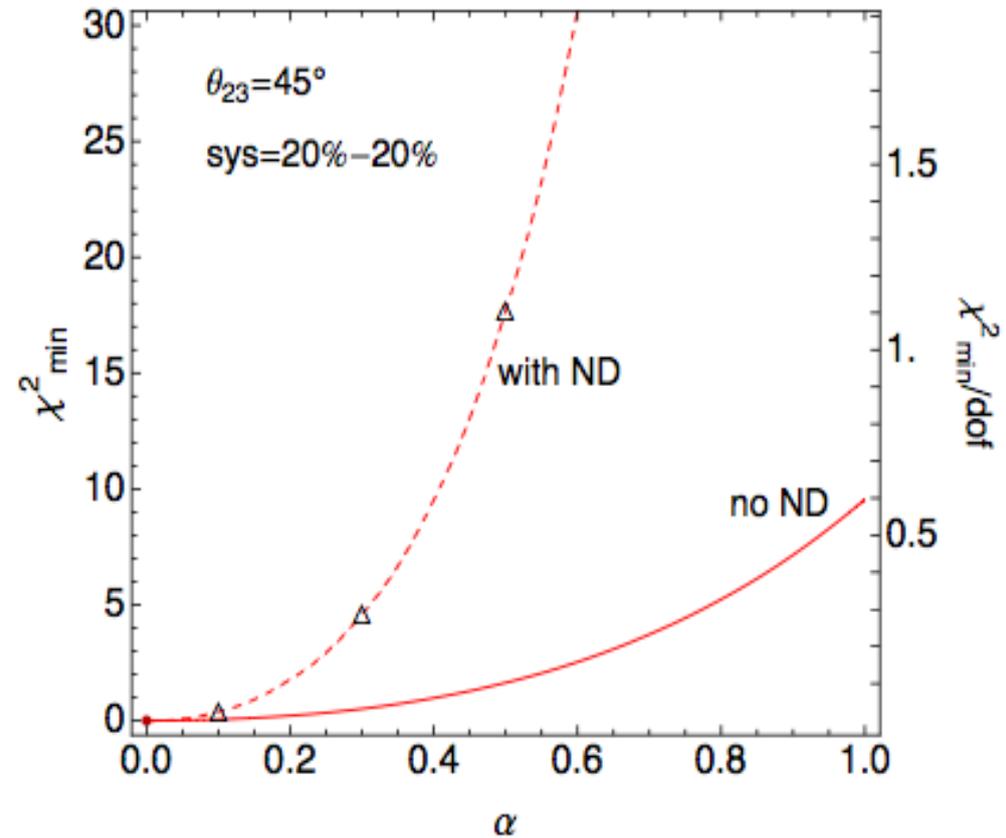
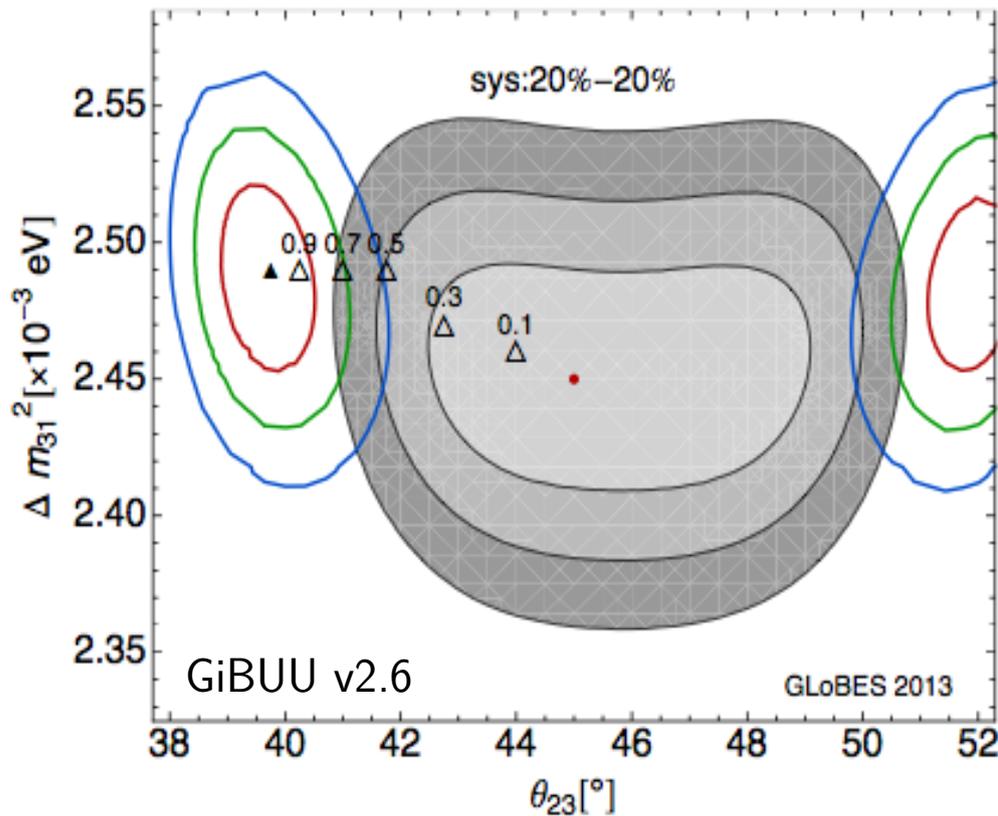
These effects can be parametrized as migration matrices from true to reconstructed energy:



Lalakulich, Mosel and Gallmeister, 1208.3678 [nucl-th]

Toy model

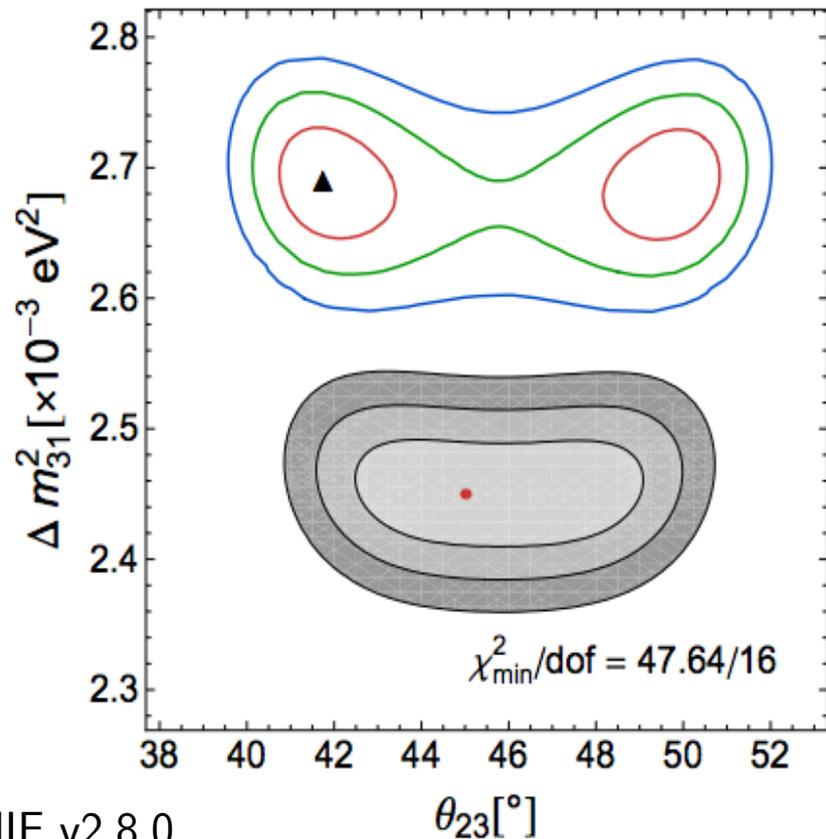
$$N_i^{\text{test}}(\alpha) = \alpha \times N_i^{\text{QE}} + (1 - \alpha) \times N_i^{\text{QE-like}}$$



Coloma and Huber, 1307.1243 [hep-ph]

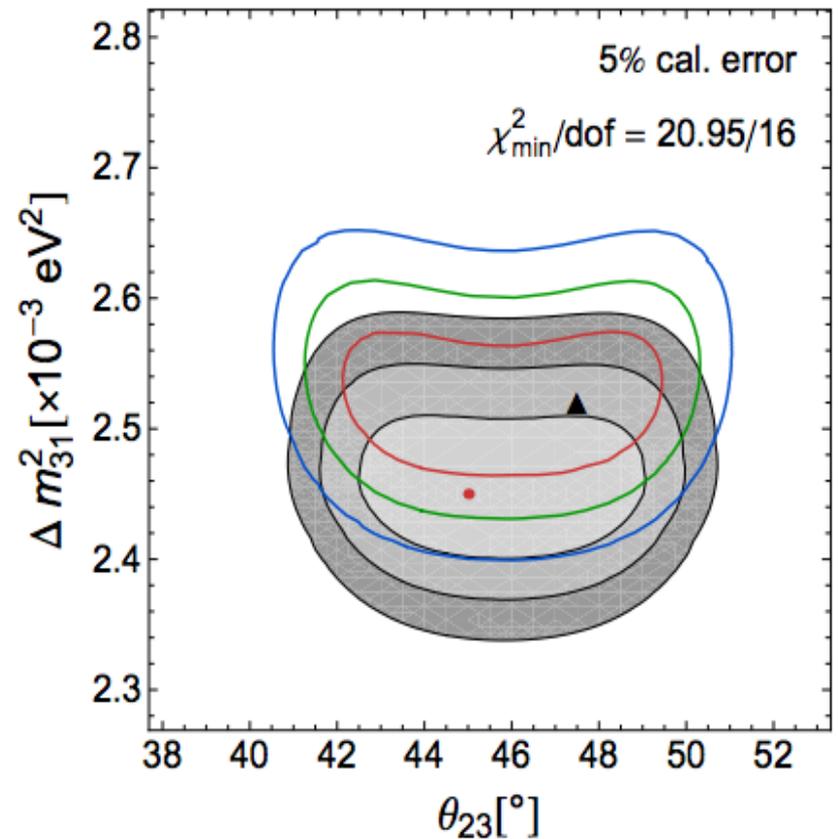
Impact of nuclear model

How large can these effects be?



GENIE v2.8.0

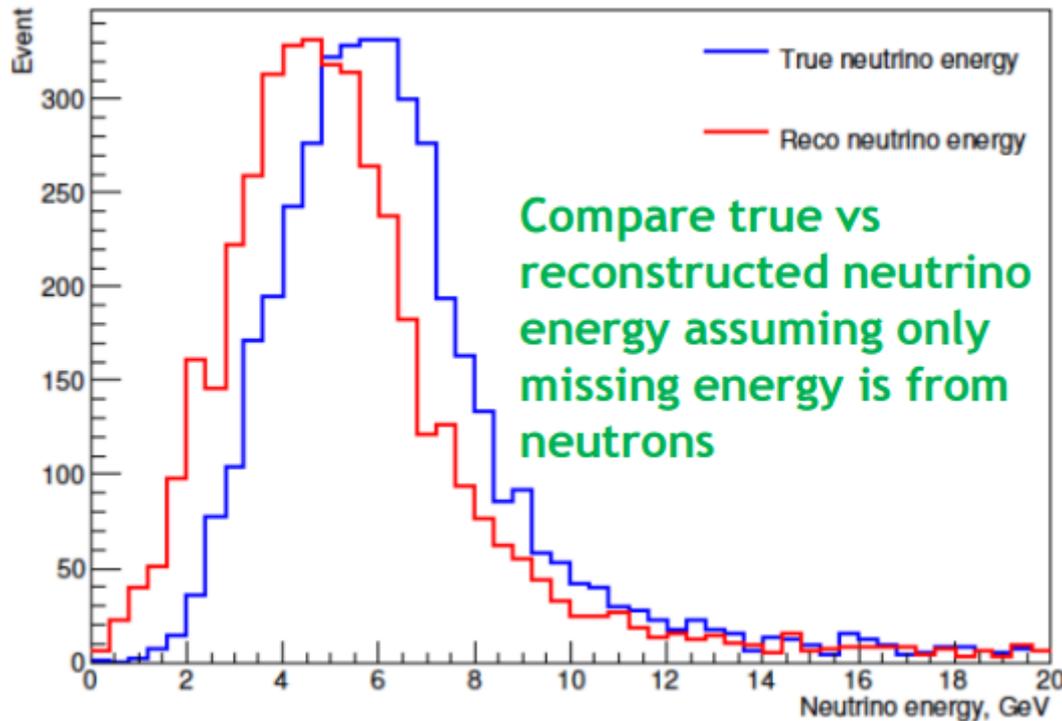
GiBUU v2.6



Coloma, Huber, Mariani and Jen, 1311.4506 [hep-ph]

Does this improve with calorimetry?

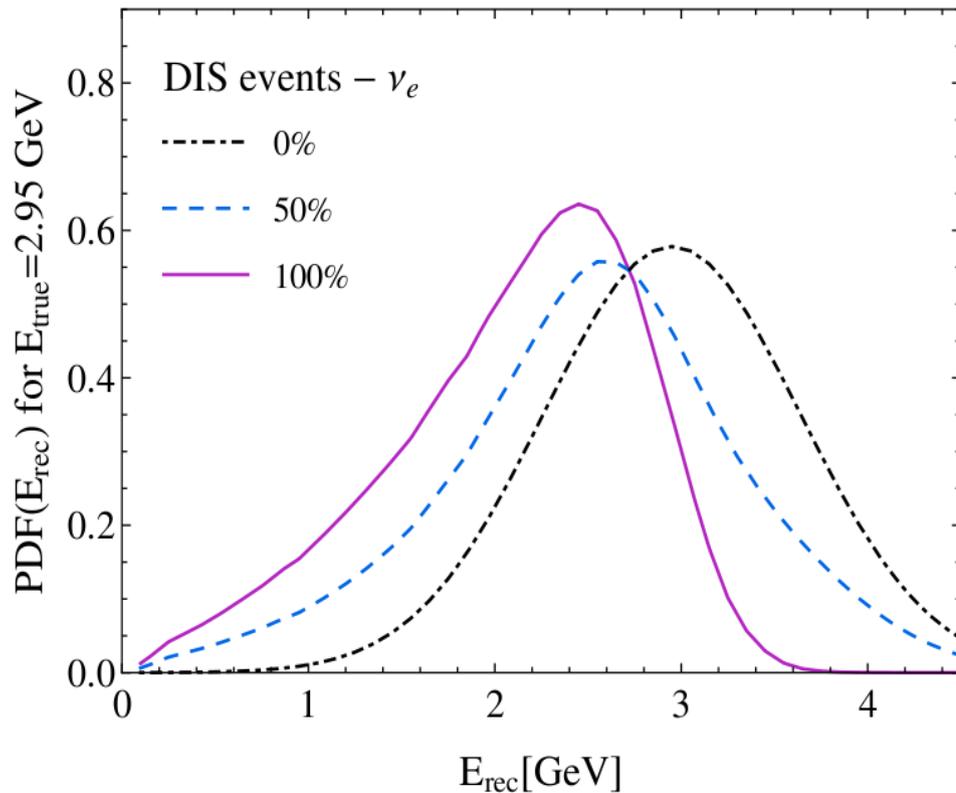
Calorimetry relies on observing all particles produced in the interaction



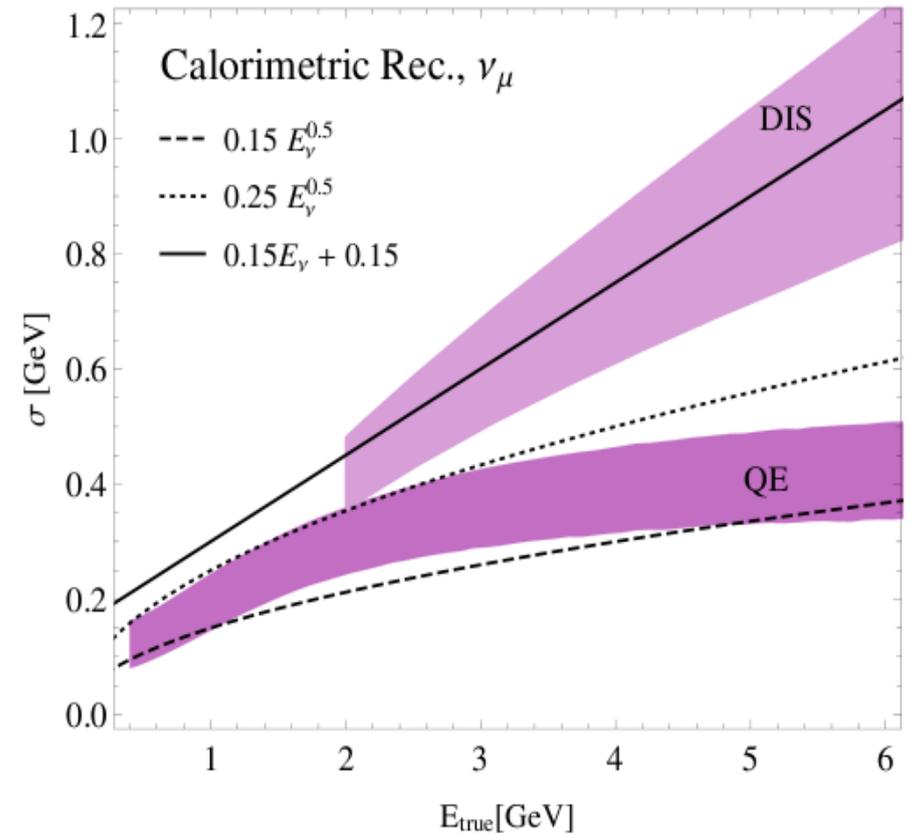
Simulation studies using NuMI medium-energy flux and GENIE (Q. Liu)

Taken from L. Whitehead's talk here at NNN

Does this improve with calorimetry?

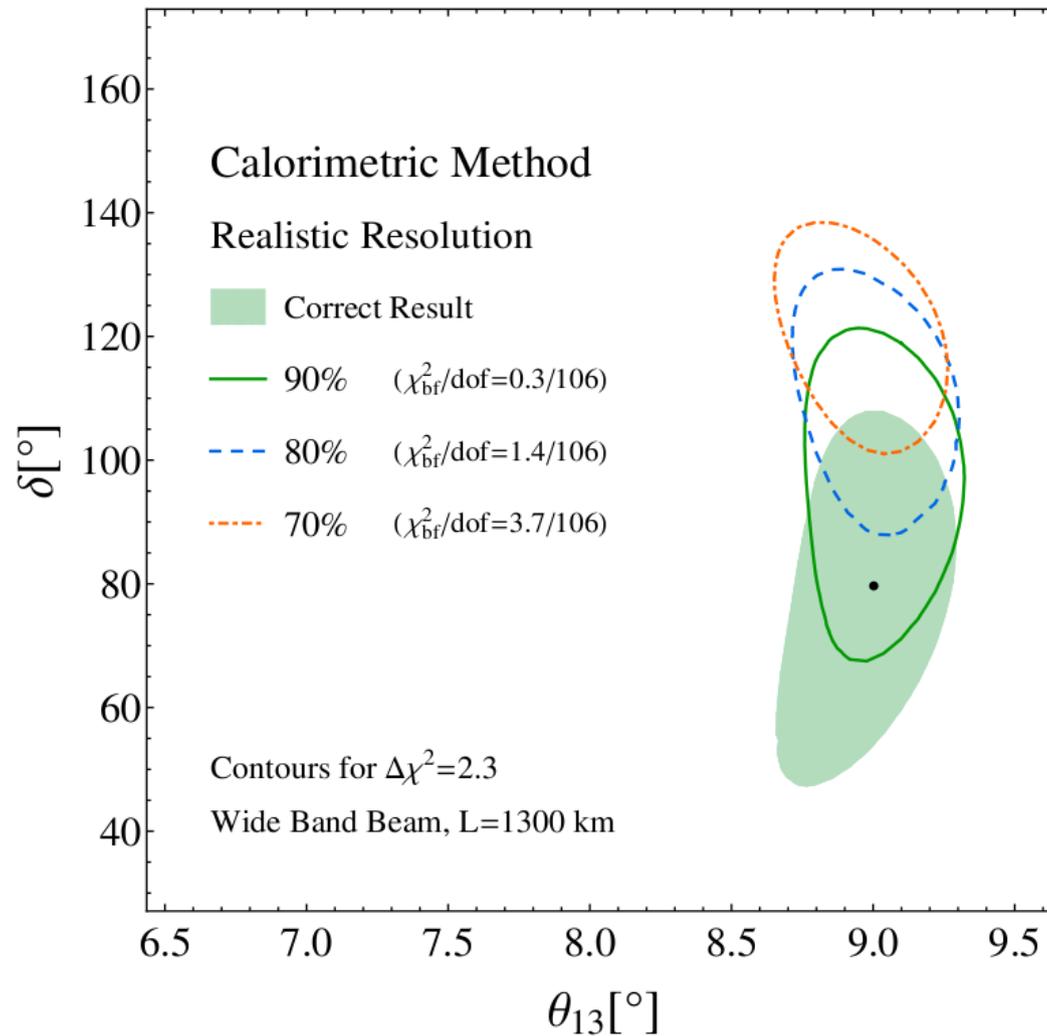


Ankowski et al, 1507.08561 [hep-ph]



Ankowski et al, 1507.08560 [hep-ph]

Does this improve with calorimetry?



Ankowski et al, 1507.08561 [hep-ph]

Summary (I/II)

- The most relevant systematics for appearance experiments are those related to cross sections. Challenges:
 - Unavailability of final flavor at the near detector
 - near-far detector extrapolation
- Systematic effects may be kept under control **under several assumptions:**
 - no flux shape uncertainties
 - no cross section shape uncertainties
 - disappearance data can be used to reduce uncertainties in appearance

challenging!

Summary (II/II)

- **Shape** uncertainties are **more dangerous**. We find:
 - Large impact on the determination of the disappearance parameters for kinematic reconstruction due to non-QE contamination of the QE sample
 - Significant bias for ΔCP from **missing energy** in calorimetric detectors, if not accurately calibrated
- Failure to include **nuclear effects** properly may induce significant bias on the oscillation parameters

Thanks!

Matrix generation details

Detection thresholds: 20 MeV for mesons and 40 MeV for protons

Detection efficiencies: 60% for pi0, 80% for other mesons, 50% for protons

Resolution details assumed:

$$\sigma(|\mathbf{k}_\mu|) = 0.02|\mathbf{k}_\mu| \quad \text{and} \quad \sigma(\theta) = 0.7^\circ,$$

$$\sigma(E_e) = 0.10E_e \quad \text{and} \quad \sigma(\theta) = 2.8^\circ$$

$$\frac{\sigma(E_{\pi^0})}{E_{\pi^0}} = \max \left\{ \frac{a_{\pi^0}}{\sqrt{E_{\pi^0}}}, \frac{b_{\pi^0}}{E_{\pi^0}} \right\} \quad \frac{\sigma(E_h)}{E_h} = \max \left\{ \frac{a_h}{\sqrt{E_h}}, b_h \right\}$$

$a_{\pi^0} = 0.107$ and $b_{\pi^0} = 0.02$, $a_h = 0.145$ and $b_h = 0.067$.

(Neutrons assumed to exit undetected)

Matrix generation details

Calorimetric reconstruction:

$$E_{\nu}^{\text{cal}} = \epsilon_n + E_{\ell} + \sum_i (E_{\mathbf{p}'_i} - M) + \sum_j E_{\mathbf{h}'_j}$$

Energy of nucleons
knocked-out

Energy of mesons
produced

Kinematic reconstruction:

$$E_{\nu}^{\text{kin}} = \frac{2(nM - \epsilon_n)E_{\ell} + W^2 - (nM - \epsilon_n)^2 - m_{\ell}^2}{2(nM - \epsilon_n - E_{\ell} + |\mathbf{k}_{\ell}| \cos \theta)}$$

Number of nucleons knocked
out of the nucleus

Invariant hadronic mass squared.

(For single-nucleon knock-out, $W^2 = M^2$)

Single nucleon
separation energy

Systematics	SB			BB			NF		
	Opt.	Def.	Cons.	Opt.	Def.	Cons.	Opt.	Def.	Cons.
Fiducial volume ND	0.2%	0.5%	1%	0.2%	0.5%	1%	0.2%	0.5%	1%
Fiducial volume FD (incl. near-far extrap.)	1%	2.5%	5%	1%	2.5%	5%	1%	2.5%	5%
Flux error signal ν	5%	7.5%	10%	1%	2%	2.5%	0.1%	0.5%	1%
Flux error background ν	10%	15%	20%	correlated			correlated		
Flux error signal $\bar{\nu}$	10%	15%	20%	1%	2%	2.5%	0.1%	0.5%	1%
Flux error background $\bar{\nu}$	20%	30%	40%	correlated			correlated		
Background uncertainty	5%	7.5%	10%	5%	7.5%	10%	10%	15%	20%
Cross secs \times eff. QE [†]	10%	15%	20%	10%	15%	20%	10%	15%	20%
Cross secs \times eff. RES [†]	10%	15%	20%	10%	15%	20%	10%	15%	20%
Cross secs \times eff. DIS [†]	5%	7.5%	10%	5%	7.5%	10%	5%	7.5%	10%
Effec. ratio ν_e/ν_μ QE [*]	3.5%	11%	–	3.5%	11%	–	–	–	–
Effec. ratio ν_e/ν_μ RES [*]	2.7%	5.4%	–	2.7%	5.4%	–	–	–	–
Effec. ratio ν_e/ν_μ DIS [*]	2.5%	5.1%	–	2.5%	5.1%	–	–	–	–
Matter density	1%	2%	5%	1%	2%	5%	1%	2%	5%

Coloma, Huber, Kopp, Winter, 1209.5973 [hep-ph]

(Theoretical constraint)

Toy model

- Neglecting all FSI and multinucleon contributions, we can compute the number of events as:

$$N_i^{QE} = \sigma_{QE}(E_i)\phi(E_i)P_{\mu\mu}(E_i)$$

- However, in practice we will observe a different distribution at the detector, given by:

$$N_i^{QE-like} = \sum_j M_{ij}^{QE} N_j^{QE} + \sum_{non-QE} \sum_j M_{ij}^{non-QE} N_j^{non-QE}$$

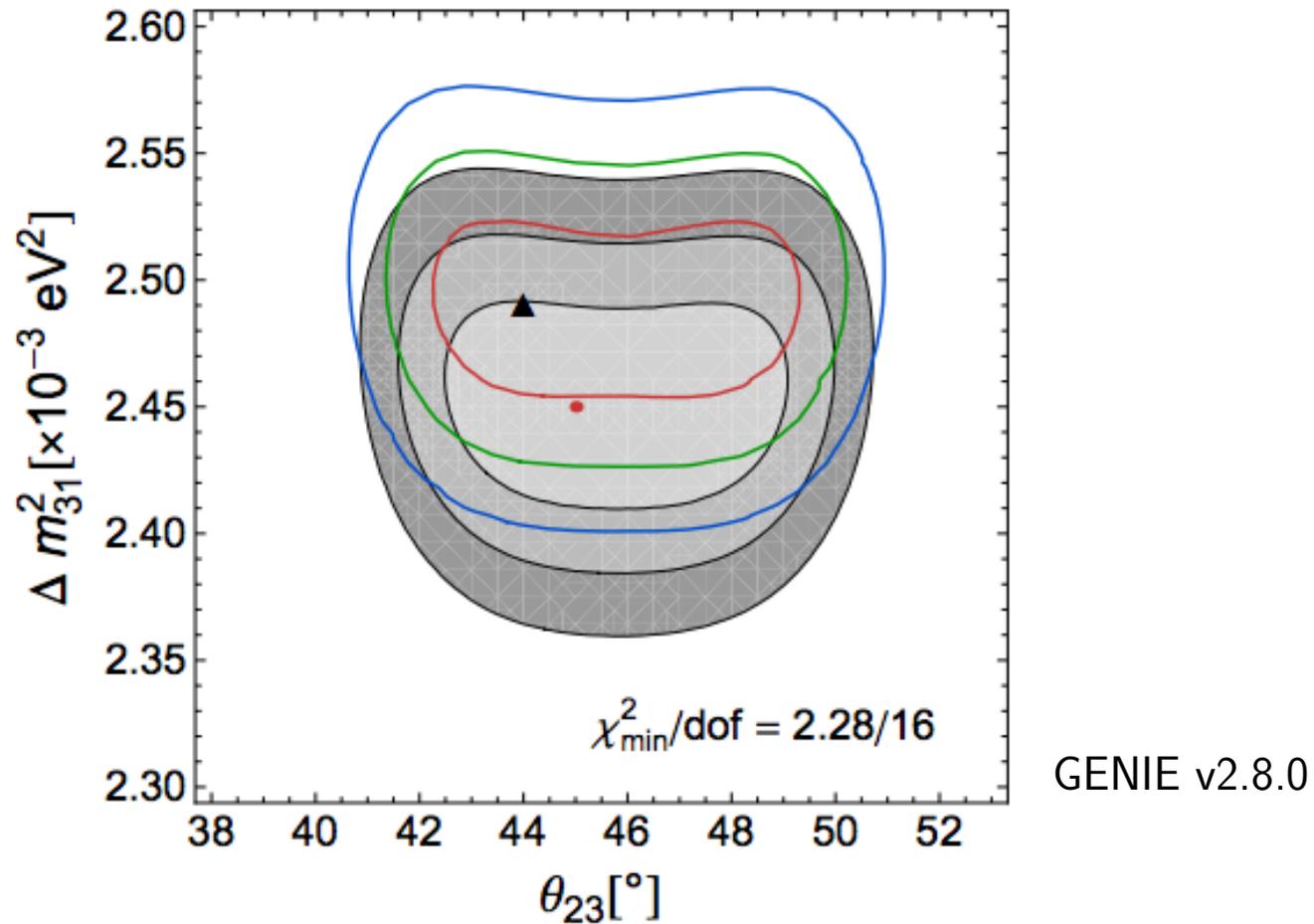
- An intermediate situation would most likely take place:

$$N_i^{test}(\alpha) = \alpha N_i^{QE} + (1 - \alpha) N_i^{QE-like}$$

Coloma and Huber, 1307.1243 [hep-ph]

Impact of target nucleus

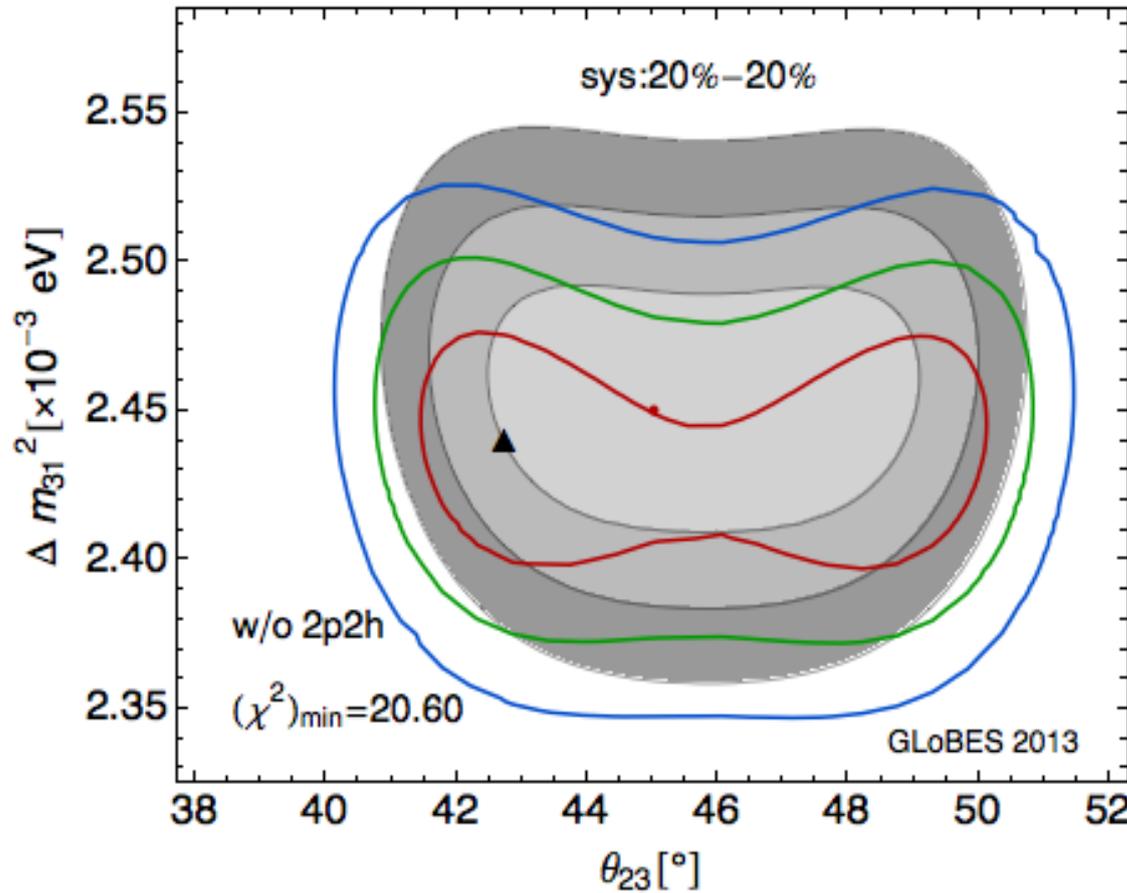
Oxygen vs Carbon:



Coloma, Huber, Mariani and Jen, 1311.4506 [hep-ph]

Impact of 2p2h

Even if we get all contributions right except 2p2h...



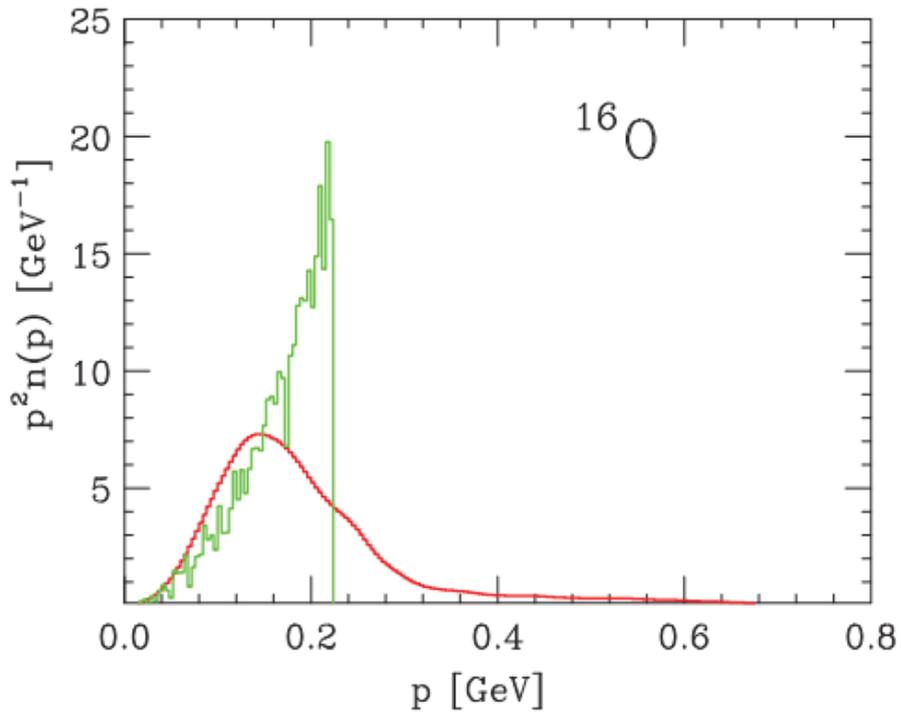
	<u>Events</u>	
	2p2h	QE-like
QE	~215	~1270
	~870	

GiBUU v2.6

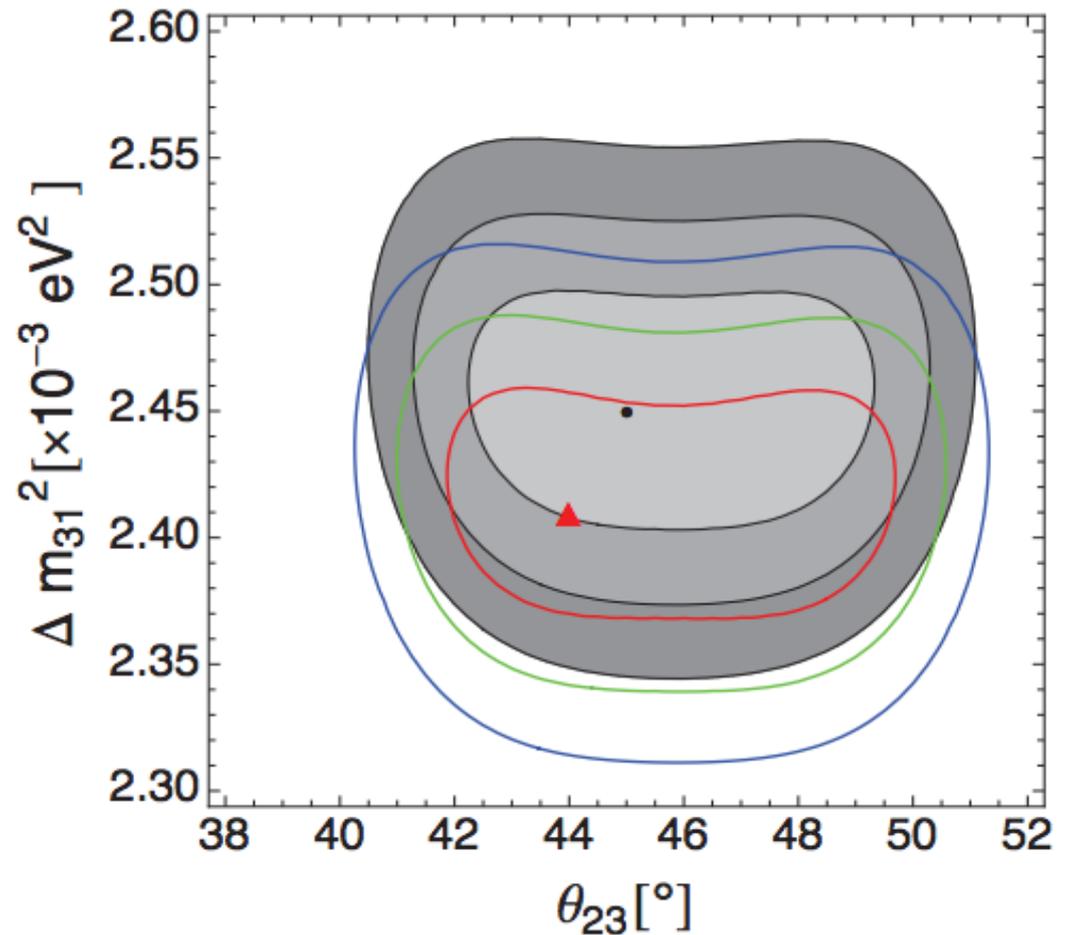
Coloma and Huber, 1307.1243 [hep-ph]

Other factors: RFGM vs SF

Nucleon momentum distribution:



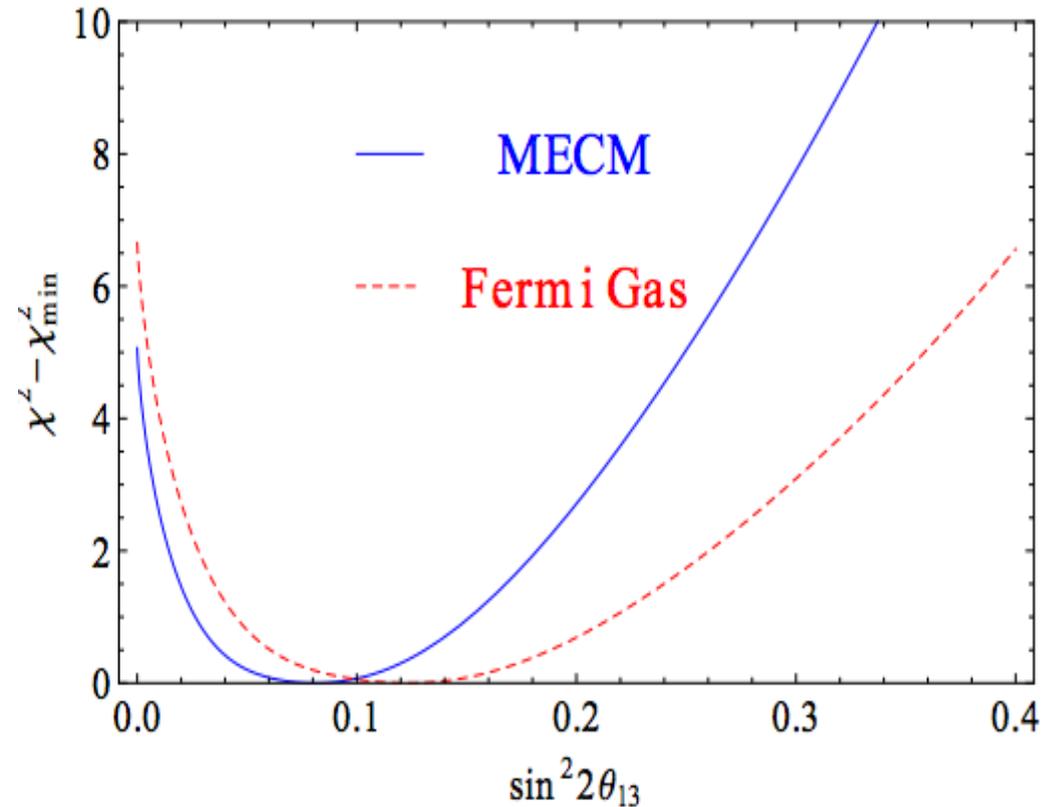
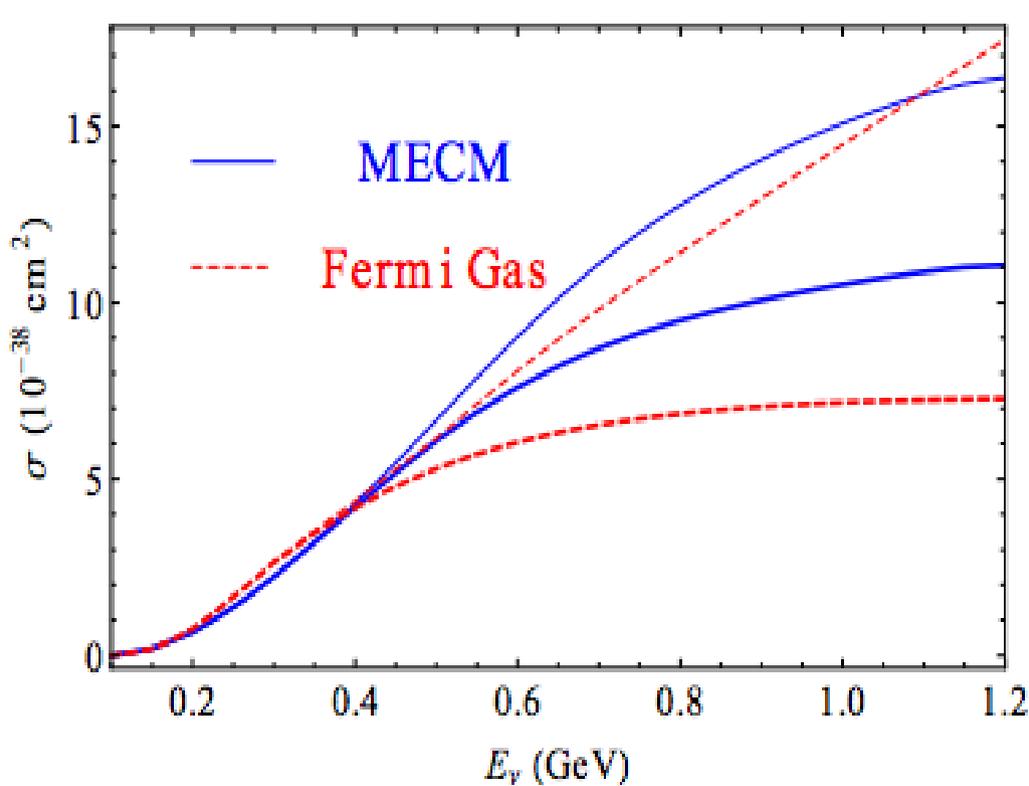
Jen et al, 1402.6651 [hep-ex]



GENIE v2.8.0 - modified

Cross section models

Impact on an analysis which reproduces T2K results in 1106.2822 [hep-ex]



Martini, Meloni, 1203.3335 [hep-ph]

MECM = model from Martini, Ericson, Chanfray, Marteau, 0910.2622 [nucl-th]